- 1. **Potential deduced from period**. The Lagrangian of a particle in d=1 is $L=mv^2/2-U(x)$, where U(-x)=U(x) and U(0)=0. U(x) has a single minimum, at x=0.
- (a) The period is T(E)=T, a constant, for all E>0. What is the potential? You may use the method of LL sec. 12.
- (b) The period is proportional to $E^{1/2}$ for all E>0. What can you say from "mechanical similarity?
- (c) Use the method of LL sec. 12 to check (b).
- **2. Planetary orbit.** Pluto's orbit has a mean distance to the sun of 39.439 au, where 1 au (astronomical unit) is 1.495×10^{11} m (distance from earth to sun.) Here "mean distance" means, not time-averaged, but spatially averaged, $r(\phi)$ averaged over ϕ . The eccentricity of its orbit is 0.2502, and inclination to the "ecliptic" is 17° . Both exceed the values for mercury, which has the largest values of the eight planets. Calculate in SI units, where $G = 6.67 \times 10^{-11} \, \text{Nm}^2/\text{kg}^2$, M (solar mass) = $1.99 \times 10^{30} \, \text{kg}$, and M (Pluto's mass) = $3 \times 10^{24} \, \text{kg}$. (a) Calculate the aphelion and perihelion distances.
- (b) Calculate the total energy (kinetic plus gravitational PE) of the sun/Pluto system (*i.e.*, neglect all other masses.)
- (c) Calculate the orbital angular momentum, and orbital period in seconds and years, of the sun/Pluto system.
- (d) If there were no other masses except the sun and Pluto, what would be the ratio of kinetic energy of the sun to kinetic energy of Pluto, in the center-of-mass frame?